## **GENERAL RELATIVITY HOMEWORK – WEEK 8**

**Exercise 1.** Consider the Schwarzschild metric induced by a central mass M:

$$ds^{2} = -\left(1 - \frac{2GM}{r}\right)dt^{2} + \frac{dr^{2}}{1 - 2GM/r} + r^{2}d\Omega^{2} .$$
 (1)

Within this spacetime, consider a particle hanging at a constant position  $(r, \theta, \phi)$ . Note that it must be using some non-gravitational force to keep it from falling inwards. Find the particle's 4-acceleration  $\alpha^{\mu} = Du^{\mu}/d\tau$  and its magnitude  $\sqrt{\alpha_{\mu}\alpha^{\mu}}$ , as a function of r. What happens as we approach r = 2GM?

**Exercise 2.** Now, consider a free-falling particle in the Schwarzschild metric, orbiting at a constant radius r. Without loss of generality, assume that the orbit is in the  $\theta = \pi/2$  plane. Find the orbit's frequency  $\omega \equiv d\phi/dt$ . Hint: use the vanishing of  $\alpha^r$  to relate  $u^t$  and  $u^{\phi}$ . What is the smallest value of r for which circular orbits are possible? Hint: the particle's worldline must be timelike!

Exercise 3. Repeat our derivation of the Schwarzschild metric, but for 3d spacetime  $(t, r, \phi)$ instead of 4d, and with circular instead of spherical symmetry. If you are weak, assume from the start that the metric is time-independent. As we did in class, find the solution to the vacuum equation  $R_{\mu\nu} = 0$ . Can you make this solution consistent with the presence of a point mass  $T^{tt} = M\delta^2(\vec{r})$  at the origin?