

GENERAL RELATIVITY HOMEWORK – WEEK 1

Exercise 1. Consider the motion $\mathbf{r}(t) = (x(t), y(t), z(t))$ of a non-relativistic object of mass m in the Earth's gravitational field, governed by the action:

$$S = m \int dt \left(\frac{1}{2}(\dot{x}^2 + \dot{y}^2 + \dot{z}^2) - gz \right), \quad (1)$$

assuming a constant free-fall acceleration g .

Find the object's trajectory, as a function of its initial position \mathbf{r}_1 at time t_1 and its final position \mathbf{r}_2 at time t_2 . Find the value of the action $S(t_2, \mathbf{r}_2; t_1, \mathbf{r}_1)$ for this trajectory. Verify that the derivatives $\partial S / \partial \mathbf{r}_2$ coincide with the momentum $\mathbf{p} = m\dot{\mathbf{r}}$ at time t_2 . Show also that $-\partial S / \partial t_2$ is the energy $E = m \left(\frac{1}{2} \dot{\mathbf{r}}^2 + gz \right)$ at t_2 .

Exercise 2. In full electromagnetism, the scalar potential $\varphi(t, \mathbf{x})$ is joined by a vector potential $\mathbf{A}(t, \mathbf{x})$, and the action reads:

$$S = \sum_n \int dt \left(\frac{m_n \dot{\mathbf{x}}_n^2}{2} + q_n (\dot{\mathbf{x}}_n \cdot \mathbf{A} - \varphi) \right) + \frac{\epsilon_0}{2} \int dt d^3\mathbf{x} (\mathbf{E}^2 - \mathbf{B}^2), \quad (2)$$

$$\text{where } \mathbf{E} = -\dot{\mathbf{A}} - \nabla\varphi \quad \text{and} \quad \mathbf{B} = \nabla \times \mathbf{A}.$$

This is the theory as it was known before Einstein: the EM field is secretly relativistic, but the particles still aren't.

1. What's the canonical momentum \mathbf{p}_n conjugate to \mathbf{x}_n ? What are the Euler-Lagrange equations of motion for \mathbf{x}_n ?
2. What are the canonical momenta $\pi_\phi, \boldsymbol{\pi}_\mathbf{A}$ conjugate to φ and \mathbf{A} ? What are the Euler-Lagrange field equations for φ and \mathbf{A} ? Express your answers in terms of \mathbf{E} and \mathbf{B} . Hint: if you're not using index notation, remember the cyclic property of $\mathbf{a} \cdot (\mathbf{b} \times \mathbf{c})$.
3. What is the energy of the particles+field system? Using the field equations and integration by parts, write it as $\frac{1}{2} \sum m_n \dot{\mathbf{x}}_n^2$ plus a spatial integral over some energy density constructed from \mathbf{E} and \mathbf{B} .
4. Now, to really test your understanding of analytical mechanics magic, derive the system's linear momentum. Again, write it as $\sum m_n \dot{\mathbf{x}}_n$ plus a spatial integral over a density constructed from \mathbf{E} and \mathbf{B} . If not using index notation, remember $(\mathbf{a} \times \mathbf{b}) \times \mathbf{c} = (\mathbf{a} \cdot \mathbf{c})\mathbf{b} - (\mathbf{b} \cdot \mathbf{c})\mathbf{a}$.