

## GENERAL RELATIVITY HOMEWORK – WEEK 1

**Exercise 1.** Consider the motion  $\mathbf{r}(t) = (x(t), y(t), z(t))$  of a non-relativistic object of mass  $m$  in the Earth's gravitational field, governed by the action:

$$S = m \int dt \left( \frac{1}{2}(\dot{x}^2 + \dot{y}^2 + \dot{z}^2) - gz \right), \quad (1)$$

assuming a constant free-fall acceleration  $g$ .

Find the object's trajectory, as a function of its initial position  $\mathbf{r}_1$  at time  $t_1$  and its final position  $\mathbf{r}_2$  at time  $t_2$ . Find the value of the action  $S(t_2, \mathbf{r}_2; t_1, \mathbf{r}_1)$  for this trajectory. Verify that the derivatives  $\partial S / \partial \mathbf{r}_2$  coincide with the momentum  $\mathbf{p} = m\dot{\mathbf{r}}$  at time  $t_2$ . Show also that  $-\partial S / \partial t_2$  is the energy  $E = m(\frac{1}{2}\dot{\mathbf{r}}^2 + gz)$  at  $t_2$ .

**Exercise 2.** In full electromagnetism, the scalar potential  $\varphi(t, \mathbf{x})$  is joined by a vector potential  $\mathbf{A}(t, \mathbf{x})$ , and the action reads:

$$S = \sum_n \int dt \left( \frac{m_n \dot{\mathbf{x}}_n^2}{2} + q_n(\dot{\mathbf{x}}_n \cdot \mathbf{A} - \varphi) \right) + \frac{\epsilon_0}{2} \int dt d^3\mathbf{x} (\mathbf{E}^2 - \mathbf{B}^2), \quad (2)$$

$$\text{where } \mathbf{E} = -\dot{\mathbf{A}} - \nabla\varphi \quad \text{and} \quad \mathbf{B} = \nabla \times \mathbf{A}.$$

This is the theory as it was known before Einstein: the EM field is secretly relativistic, but the particles still aren't.

1. What's the canonical momentum  $\mathbf{p}_n$  conjugate to  $\mathbf{x}_n$ ? What are the Euler-Lagrange equations of motion for  $\mathbf{x}_n$ ?
2. What are the canonical momenta  $\pi_\phi, \boldsymbol{\pi}_\mathbf{A}$  conjugate to  $\varphi$  and  $\mathbf{A}$ ? What are the Euler-Lagrange field equations for  $\varphi$  and  $\mathbf{A}$ ? Express your answers in terms of  $\mathbf{E}$  and  $\mathbf{B}$ . Hint: if you're not using index notation, remember the cyclic property of  $\mathbf{a} \cdot (\mathbf{b} \times \mathbf{c})$ .
3. What is the energy of the particles+field system? Using the field equations and integration by parts, write it as  $\frac{1}{2} \sum m_n \dot{\mathbf{x}}_n^2$  plus a spatial integral over some energy density constructed from  $\mathbf{E}$  and  $\mathbf{B}$ .
4. Now, to really test your understanding of analytical mechanics magic, derive the system's linear momentum. Again, write it as  $\sum m_n \dot{\mathbf{x}}_n$  plus a spatial integral over a density constructed from  $\mathbf{E}$  and  $\mathbf{B}$ . If not using index notation, remember  $(\mathbf{a} \times \mathbf{b}) \times \mathbf{c} = (\mathbf{a} \cdot \mathbf{c})\mathbf{b} - (\mathbf{b} \cdot \mathbf{c})\mathbf{a}$ .