

Exploring complex saddles and geometries through holography.

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- ▶ When studying the path integral for quantum gravity, often we consider the complex analytic continuation of the metric, however some interesting puzzles can arise.
- ▶ A classic example comes from complexifying S^{d+1} : [Witten 2021].

$$ds^2 = \ell^2 \left[\left(\frac{d\theta(u)}{du} \right)^2 du^2 + \cos^2 \theta(u) d\Omega_d^2 \right]$$

When $\theta(u) = u$, $0 \leq u \leq \pi$, we have S^{d+1} , however when $\theta(u) = iu$, $-\infty \leq u \leq +\infty$, we have Lorentzian dS_{d+1} .

- ▶ If the universe started from nothing, i.e. $\cos^2 \theta = 0$, we can have:

$$\theta = \left(n + \frac{1}{2} \right) \pi, \quad n \in \mathbb{Z}$$

and evolve into $\theta = iu$ as $u \rightarrow +\infty$. We thus have a family of complex geometries with initial conditions labeled by $n \in \mathbb{Z}$.

- ▶ More precisely, we complexified the action $\Psi \sim \exp[S_{(n)} + i\mathcal{I}(u)]$, where real part $S_{(n)} = (n + \frac{1}{2})S$, where S is the de Sitter entropy. This implies we can increase the amplitude as $n \rightarrow \infty$.
- ▶ In principle we need to sum over all $n \in \mathbb{Z}$ when evaluating the path integral, however the proposal by Hartle-Hawking only needs $\theta(0) = \pm \frac{\pi}{2}$ or $n = 0, -1$.
- ▶ Kontsevich-Segal: For \mathcal{M}^D with complex metric g , consider a p -form $A^{(p)}$ with field strength $dA^{(p)}$ and the action with $q = p + 1$:

$$I_q = \frac{1}{2q!} \int_M d^D x \sqrt{\det g} g^{i_1 j_1} \dots g^{i_q j_q} F_{i_1 i_2 \dots i_q} F_{j_1 j_2 \dots j_q}$$

the allowable g is such that:

$$\text{Re} \left(\sqrt{\det g} g^{i_1 j_1} \dots g^{i_q j_q} F_{i_1 i_2 \dots i_q} F_{j_1 j_2 \dots j_q} \right) > 0, \quad 0 \leq q \leq D,$$

for all non-zero $p+1$ form $F^{(p+1)} = dA^{(p)}$. [\[Witten 2021\]](#).

Holography for AdS and dS spacetimes

Some ingredients for building AdS_{d+1}/CFT_d correspondence from the bottom up.

- ▶ Starting with the Poincaré coordinates: (y, \vec{x}) , $y \geq 0$ of AdS_{d+1}:

$$ds^2 = \ell_{\text{AdS}}^2 \frac{dy^2 + d\vec{x}^2}{y^2}$$

- ▶ The bulk Scalar/Tensor Fields in AdS_{d+1}, $\phi^{\text{AdS}}(y, \vec{x})/\sigma_{i_1 \dots i_s}^{\text{AdS}}(y, \vec{x})$ with mass m and spin s :

$$\ell_{\text{AdS}}^2 m^2 = -(\Delta_+ \Delta_- + s), \quad \Delta_- = d - \Delta_+$$

approaching AdS boundary at $y \rightarrow 0^+$, these bulk fields become:

$$\phi^{\text{AdS}}(y, \vec{x}) \sim \phi_+^{\text{AdS}}(\vec{x}) y^{\Delta_+} + \phi_-^{\text{AdS}}(\vec{x}) y^{\Delta_-},$$

$$\sigma_{i_1 \dots i_s}^{\text{AdS}}(y, \vec{x}) \sim \sigma_{i_1 \dots i_s}^{+, \text{AdS}}(\vec{x}) y^{\Delta_+ - s} + \sigma_{i_1 \dots i_s}^{-, \text{AdS}}(\vec{x}) y^{\Delta_- - s}.$$

Similar proposal was made for dS_{d+1}/CFT_d :

Strominger 2001, Maldacena 2002, Anninos-Hartman-Strominger 2011.

- ▶ Starting with the Poincaré coordinates: (η, \vec{x}) , $\eta \leq 0$ for dS_{d+1} :

$$ds^2 = \ell^2 \frac{-d\eta^2 + d\vec{x}^2}{\eta^2}$$

- ▶ The bulk Scalar/Tensor Fields in dS_{d+1} , $\phi(y, \vec{x})/\sigma_{i_1 \dots i_s}(y, \vec{x})$ with masses:

$$\ell^2 m^2 = \Delta_+ \Delta_- + s, \quad \Delta_- = d - \Delta_+$$

as $\eta \rightarrow 0^-$, approaching future infinity, the bulk fields become:

$$\begin{aligned} \phi(\eta, \vec{x}) &\sim \phi_+(\vec{x}) (-\eta)^{\Delta_+} + \phi_-(\vec{x}) (-\eta)^{\Delta_-} \\ \sigma_{i_1 \dots i_s}(\eta, \vec{x}) &\sim \sigma_{i_1 \dots i_s}^+(\vec{x}) (-\eta)^{\Delta_+ - s} + \sigma_{i_1 \dots i_s}^-(\vec{x}) (-\eta)^{\Delta_- - s} \end{aligned}$$

We can also choose $\eta \geq 0$, the boundary is at past infinity $\eta \rightarrow 0^+$.

- ▶ If we relate $\text{AdS}_{d+1}/\text{dS}_{d+1}$ metrics via analytic continuation:

$$y = -i\eta, \quad \ell_{\text{AdS}} = -i\ell.$$

From the boundary couplings between the bulk fields and the dual CFT operators in AdS/CFT correspondence:

$$\ell_{\text{AdS}}^d \int d^d \vec{x} \phi_{\pm}^{\text{AdS}} \mathcal{O}_{\text{AdS}}^{\pm}, \quad \ell_{\text{AdS}}^d \int d^d \vec{x} \sigma_{i_1 \dots i_s}^{\pm, \text{AdS}} J_{\pm, \text{AdS}}^{i_1 \dots i_s}.$$

- ▶ For dS/CFT correspondence, the coupling is via the wave functional:

$$\Psi[\psi_j^0] = \left\langle \exp \left(\ell^d \int d^d \vec{x} \psi_j^0 \mathcal{O}^j \right) \right\rangle.$$

Identifying the two couplings leads us to propose:

$$\phi_{\pm} = e^{i\frac{\pi}{2}\Delta_{\pm}} \phi_{\pm}^{\text{AdS}} \quad \Leftrightarrow \quad \mathcal{O}_{\pm} = e^{i\frac{\pi}{2}(d-\Delta_{\pm})} \mathcal{O}_{\pm}^{\text{AdS}}.$$

where the identifications come from the asymptotic behavior.

- ▶ While for the AdS/dS bulk tensor fields, we have the identifications:

$$\sigma_{i_1 \dots i_s}^{\pm}(\vec{x}) = e^{i\frac{\pi}{2}\Delta_{\pm}} \sigma_{i_1 \dots i_s}^{\pm, \text{AdS}}(\vec{x}), \quad J_{i_1 \dots i_s}^{\pm}(\vec{x}) = e^{i\frac{\pi}{2}(d-\Delta_{\pm})} J_{i_1 \dots i_s}^{\pm, \text{AdS}}(\vec{x}).$$

after including raising/lowering of tensor indices. For spin s conserved current: $\Delta_+ = s + d - 2$, this yields:

$$J_{i_1 \dots i_s}^+(\vec{x}) = e^{i\frac{\pi}{2}(2-s)} J_{i_1 \dots i_s}^{+, \text{AdS}}(\vec{x})$$

which implies the definition of boundary energy momentum tensor remains invariant. The asymptotic symmetry which is given by the energy momentum tensor, is preserved.

- ▶ However things become more complicated for correlation functions, as different analytic continuations can be applied to different fields as they are all space-like separated, which account for time (anti-) ordering in the in-in formalism.

An Explicit Higher Spin dS_3/CFT_2 Correspondence

- ▶ Starting with pure AdS_3 gravity, we consider $SL(2, \mathbb{R})^2$ Chern-Simons gauge theory:

$$S = S_{CS}[A] - S_{CS}[\tilde{A}], \quad S_{CS}[A] = \frac{k}{4\pi} \int \text{tr} \left(A \wedge dA + \frac{2}{3} A \wedge A \wedge A \right)$$

and CS level k is related to G_N and ℓ_{AdS} via:

$$k = \frac{\ell_{AdS}}{4G_N}.$$

- ▶ The independent gauge fields A and \tilde{A} take values in each copy of $SL(2, \mathbb{R})$ algebra: $[L_m, L_n] = (m - n)L_{m+n}$, $m, n = 0, \pm 1$

$$A = e^{-\rho L_0} a e^{\rho L_0} + L_0 d\rho, \quad \tilde{A} = e^{\rho L_0} \tilde{a} e^{-\rho L_0} - L_0 d\rho$$

$$a = a_+(x^+) dx^+, \quad \tilde{a} = \tilde{a}_-(x^-) dx^-$$

where $a_+(x^+)$, $\tilde{a}_-(x^-)$ are arbitrary functions of $x^\pm = t \pm \phi$ with $\phi \sim \phi + 2\pi$.

- For the AdS₃ BTZ black hole, it is given by:

$$a_+(x^+) = L_1 - \frac{2\pi\mathcal{L}^{\text{AdS}}}{k}L_{-1}, \quad \tilde{a}_-(x^-) = -L_{-1} + \frac{2\pi\mathcal{L}^{\text{AdS}}}{k}L_1,$$

$$\mathcal{L}_{\text{AdS}} = \frac{\ell^{\text{AdS}}r_+^2}{32\pi G_N} = \frac{kr_+^2}{8\pi}.$$

which yield the following metric:

$$\begin{aligned} \ell_{\text{AdS}}^{-2}ds^2 = & d\rho^2 - \left(e^\rho - \frac{2\pi\mathcal{L}^{\text{AdS}}}{k}e^{-\rho} \right) \left(e^\rho - \frac{2\pi\mathcal{L}^{\text{AdS}}}{k}e^{-\rho} \right) dt^2 \\ & + \left(e^\rho + \frac{2\pi\mathcal{L}^{\text{AdS}}}{k}e^{-\rho} \right) \left(e^\rho + \frac{2\pi\mathcal{L}^{\text{AdS}}}{k}e^{-\rho} \right) d\phi^2 \end{aligned}$$

and we can recover the usual BTZ black hole by coordinate change:

$$r = e^\rho + \frac{2\pi\mathcal{L}^{\text{AdS}}}{k}e^{-\rho}$$

- ▶ Taking $t \rightarrow it_E$, the absence of conical singularity at the horizon r_+ demands following periodicity:

$$t_E \sim t_E + \beta^{\text{AdS}}, \quad \beta^{\text{AdS}} = \frac{2\pi}{r_+}$$

- ▶ The Chern-Simons gauge configurations can be classified by gauge invariant Wilson loop:

$$\mathcal{P}e^{\oint A} = \mathcal{P}e^{\oint dt_E A_{t_E}} = e^{-i\theta L_0} e^{\Omega} e^{i\theta L_0}.$$

The BTZ black hole corresponds to the eigenvalues of holonomy Ω equals $(+\pi, -\pi)$, however if we consider large gauge transformations, other values $(2\pi(n + \frac{1}{2}), -2\pi(n + \frac{1}{2}))$, $n \in \mathbb{Z}^+$ are also allowed.

We proposed an explicit dS_3/CFT_2 correspondence by suitable analytic continuation of AdS_3/CFT_2 one.

[Hikida, Nishioka, Takayanagi, Taki, 2022], [Chen, Hikida 2022], [Chen, Chen, Hikida 2022]

- ▶ We analytically continue the CS-theory to $G = SL(2, \mathbb{C})^2$:

$$S = S_{CS}[A] - S_{CS}[\bar{A}], \quad S_{CS}[A] = -\frac{\kappa}{4\pi} \int \text{tr} \left(A \wedge dA + \frac{2}{3} A \wedge A \wedge A \right)$$

$$\kappa = \frac{\ell}{4G_N}.$$

We have set $\ell_{AdS} \rightarrow i\ell$ such that $k \rightarrow i\kappa$. Also $e^{-\rho} = e^{-(\tilde{\rho} + i\frac{\pi}{2})}$.

- ▶ We need to impose the complex conjugation as: $(L_0)^* = -L_0$, $(L_{\pm})^* = L_{\mp}$. The gauge fields are now expressed as:

$$A = e^{-(\tilde{\rho} + \pi i/2)L_0} a e^{(\tilde{\rho} + \pi i/2)L_0} + L_0 d\tilde{\rho}, \quad \bar{A} = e^{(\tilde{\rho} + \pi i/2)L_0} \bar{a} e^{-(\tilde{\rho} + \pi i/2)L_0} - L_0 d\tilde{\rho}$$

$$a = a_+(x^+) dx^+, \quad \bar{a} = \bar{a}_-(x^-) dx^-.$$

where $x^{\pm} = it \pm \phi$, $\phi \sim \phi + 2\pi$.

- ▶ To obtain the dS_3 analogue of BTZ black hole, we can consider the following configuration:

$$a_+(x^+) = L_1 + \frac{2\pi\mathcal{L}}{\kappa}L_{-1}, \quad \bar{a}_-(x^-) = -L_{-1} - \frac{2\pi\mathcal{L}}{\kappa}L_1,$$

- ▶ Following the coordinate shift: $\tilde{\rho} \rightarrow \tilde{\rho} + \log \sqrt{\mathcal{L}/\kappa}$ and continuation $\tilde{\rho} = i\theta$, we have:

$$\ell^{-2}ds^2 = d\theta^2 - \frac{8\pi\mathcal{L}}{\kappa} \sin^2 \theta dt^2 + \frac{8\pi\mathcal{L}}{\kappa} \cos^2 \theta d\phi^2.$$

which can be mapped into dS_3 BTZ black hole metric via the coordinate and parameter transformations:

$$r = \sqrt{\frac{8\pi\mathcal{L}}{\kappa}} \cos \theta, \quad \mathcal{L} = \frac{\ell r_+^2}{32\pi G_N} = \frac{\kappa r_+^2}{8\pi}$$

- ▶ This allows us to obtain asymptotically dS₃ BTZ black hole geometry:

$$ds^2 = \ell^2 \left[-(r_+^2 - r^2)dt^2 + \frac{1}{r_+^2 - r^2}dr^2 + r^2d\phi^2 \right]$$

Here $\phi \sim \phi + 2\pi$ and the horizon is at $r_+ = \sqrt{1 - 8G_N E}$. Under $t \rightarrow -it_E$, the absence of conical singularity at horizon needs $t_E \sim t_E + 2\pi/r_+$. The Gibbons-Hawking entropy of dS₃ BTZ black hole is thus:

$$S_{\text{GH}} = \frac{2\pi\ell r_+}{4G_N} = \frac{\pi\ell\sqrt{1 - 8G_N E}}{2G_N}$$

- ▶ We can define the holonomy matrix for the gauge field A along the compactified time cycle as:

$$\mathcal{P}e^{\oint A} = \mathcal{P}e^{\oint dt E A_+} = e^{-(i\theta + \pi i/2)L_0} e^{\Omega} e^{(i\theta + \pi i/2)L_0}$$

the eigenvalues for Ω are $(i\pi, -i\pi)$.

- ▶ Applying large gauge transformation, Ω now takes the values $(2\pi(n + \frac{1}{2}), -2\pi(n + \frac{1}{2}))$, $n \in \mathbb{Z}^+$, given by:

$$a = -\sqrt{\frac{2\pi\mathcal{L}}{\kappa}}\sigma_1(d\phi + (2n + 1)dt_E), \quad \bar{a} = -\sqrt{\frac{2\pi\mathcal{L}}{\kappa}}\sigma_1(d\phi - (2n + 1)dt_E)$$

where σ_1 is a Pauli matrix. They generate the following metric:

$$ds^2 = \ell^2 \left[d\theta^2 + \frac{8\pi(2n + 1)^2\mathcal{L}}{\kappa} \sin^2\theta dt_E^2 + \frac{8\pi\mathcal{L}}{\kappa} \cos^2\theta d\phi^2 \right].$$

if we set $r = \sqrt{1 - 8G_N E} \cos\theta$ and $\frac{8\pi\mathcal{L}}{\kappa} = 1 - 8G_N E$.

- ▶ Each saddle point contribute to the Chern-Simons path integral with the action after Euclideanization:

$$-S (\equiv S_{\text{GH}}^{(n)}) = 4\pi(2n + 1)\sqrt{2\pi\kappa\mathcal{L}} = (2n + 1)\frac{\pi\ell\sqrt{1 - 8G_N E}}{2G_N}$$

these equivalent saddles need to be summed over, however they may lead to over-counting.

Using the explicit dS_3/CFT_2 , we can select the correct saddle points from dual CFT. [Chen, Hikida, Taki, Uetoko, 2023]

- ▶ We start with the regularized action for Liouville theory:

$$S_L = \frac{1}{4\pi} \int_D d^2\sigma [\partial_a\phi\partial_a\phi + 4\pi\mu e^{2b\phi}] + \frac{Q}{\pi} \oint_{\partial D} \phi d\theta + 2Q^2 \ln R$$

- ▶ The vertex operator considered take the form:

$$V_\alpha = e^{2\alpha\phi}, \quad h = \bar{h} = \alpha(Q - \alpha)$$

- ▶ While the background charge, central charge and coupling are related via:

$$c = 1 + 6Q^2 = 1 + 6(b + b^{-1})^2$$

and it can be related gravitational constant via:

$$c(\equiv ic^{(g)}) = i \cdot 6\kappa = i \frac{3\ell}{2G_N}$$

the classical limit $G_N \rightarrow 0$ corresponds to large c limit.

- ▶ For dS_3 BTZ black hole, it corresponds to the insertion of two heavy operators in such a limit:

$$b^{-2} = \frac{ic^{(g)}}{6} - \frac{13}{6} + \mathcal{O}((c^{(g)})^{-1}), \quad b \rightarrow 0$$

We also scale $\alpha = \eta/b$ and $\phi_c = 2b\phi$, $\lambda = \pi\mu b^2$ to obtain:

$$b^2 S_L = \frac{1}{16\pi} \int_D d^2\sigma [\partial_a \phi_c \partial_a \phi_c + 16\lambda e^{\phi_c}] + \frac{1}{2\pi} \oint_{\partial D} \phi_c d\theta + 2 \ln R + \mathcal{O}(b^2)$$

Here η is kept fixed such that $2h = 2\alpha(Q - \alpha) = i\ell E$ and $1 - 2\eta = \sqrt{1 - 8G_N E}$. Need $0 < \eta < \frac{1}{2}$ for existence.

- ▶ The path integral for the two point function reduces to

$$\langle V_\alpha(z_1) V_\alpha(z_2) \rangle \equiv \int \mathcal{D}\phi_c e^{-S_L} \exp(b^{-1}\alpha(\phi_c(z_1) + \phi_c(z_2)))$$

i. e. the operator insertions become δ -sources in this heavy-limit.

- ▶ From the back reacted action, we can deduce the E.O.M for ϕ_c :

$$\partial\bar{\partial}\phi_c = 2\lambda e^{\phi_c} - 2\pi\eta[\delta^{(2)}(z - z_1) + \delta^{(2)}(z - z_2)], \quad \lambda \equiv \pi\mu b^2$$

where near $z_{1,2}$, we set $\phi_c(z) \sim -4\eta|z - z_{1,2}|$, creating conical deficits on the CFT world sheet metric. We can easily generalize to higher point correlation functions.

- ▶ We have multiple allowed solutions $\phi_{c(n)} = \phi_{c(0)} + 2\pi n$, which yields the on-shell action:

$$\begin{aligned} b^2\tilde{S}_L^{(n)} &= 2\pi i(n + 1/2)(1 - 2\eta) + (2\eta - 1)\ln \lambda \\ &+ 4(\eta - \eta^2)\ln|z_{12}| + 2[(1 - 2\eta)\ln(1 - 2\eta) - (1 - 2\eta)]. \end{aligned}$$

where $\tilde{S}_L^{(n)}$ is the modified action including the back reaction and in principle need to sum over all solutions labeled by n , which precisely corresponds to the CS monodromy.

- ▶ Happily the exact expression for the Liouville two point function is known: [DOZZ 1994, 1995].

$$\langle V_\alpha(z_1)V_\alpha(z_2)\rangle = |z_{12}|^{-4\alpha(Q-\alpha)} \frac{2\pi}{b^2} [\pi\mu\gamma(b^2)]^{(Q-2\alpha)/b} \gamma\left(\frac{2\alpha}{b} - 1 - \frac{1}{b^2}\right) \gamma(2b\alpha - b^2) \delta(0)$$

- ▶ Taking the semi-classical limit of the exact two point function: [Harlow, Maltz, Witten 2011].

$$\langle V_\alpha(z_1)V_\alpha(z_2)\rangle \sim \delta(0) |z_{12}|^{-4\eta(1-\eta)/b^2} \lambda^{(1-2\eta)/b^2} \times \left(e^{-\pi i(1-2\eta)/b^2} - e^{\pi i(1-2\eta)/b^2} \right) \exp \left\{ -\frac{2}{b^2} [(1-2\eta) \ln(1-2\eta) - (1-2\eta)] \right\}.$$

and $\delta(0)$ comes from setting $\alpha = \alpha'$. In obtaining this limit, it is crucial that $\text{Re}(b^{-2}) < 0$, i. e. consistent with our earlier relation between b and $ic^{(g)}$ for dual CFT to de Sitter. This result can only be reproduced by $\tilde{S}_L^{(n)}$ with $n = 0, -1$, leading us to correct saddles.

- ▶ Taking the modulus of the two point function, where the $|z_{12}|$ dependence now cancels out, as $1/b^2 \sim ic^{(g)}/6$ is purely imaginary, we have:

$$|\langle V_\alpha(z_1)V_\alpha(z_2)\rangle| \sim \left| e^{\frac{\pi c^{(g)}}{6}\sqrt{1-8G_N E}} - e^{-\frac{\pi c^{(g)}}{6}\sqrt{1-8G_N E}} \right|$$

the leading order contribution precisely reproduces the Gibbons-Hawking entropy S_{GH} of the corresponding dS black holes.

- ▶ This also implies that identification of the phase of the two point function:

$$\langle V_\alpha(z_1)V_\alpha(z_2)\rangle \sim \Psi \sim \exp\left(\frac{S_{\text{GH}}}{2} + i\mathcal{I}\right) \implies \mathcal{I} = \frac{c^{(g)}}{6} \log \lambda$$

such that λ is manifestly real. Consistent with the earlier CFT computation for the phase in [\[Hikida, Nishioka, Takayanagi, Taki, 2022\]](#)

It is also interesting to employ similar strategy to investigate the bulk geometries dual to higher point correlation functions of heavy operators.

- ▶ For three point function, we have: [\[DOZZ 1994, 1995\]](#)

$$\langle V_{\alpha_1}(z_1, \bar{z}_1) V_{\alpha_2}(z_2, \bar{z}_2) V_{\alpha_3}(z_3, \bar{z}_3) \rangle = \frac{C(\alpha_1, \alpha_2, \alpha_3)}{|z_{12}|^{2(h_1+h_2-h_3)} |z_{13}|^{2(h_1+h_3-h_2)} |z_{23}|^{2(h_2+h_3-h_1)}}$$

$$C(\alpha_1, \alpha_2, \alpha_3) = \left[\lambda \gamma(b^2) b^{-2b^2} \right]^{(Q - \sum_i \alpha_i)/b}$$

$$\times \frac{\Upsilon'_b(0) \Upsilon_b(2\alpha_1) \Upsilon_b(2\alpha_2) \Upsilon_b(2\alpha_3)}{\Upsilon_b(\sum_i \alpha_i - Q) \Upsilon_b(\alpha_1 + \alpha_2 - \alpha_3) \Upsilon_b(\alpha_2 + \alpha_3 - \alpha_1) \Upsilon_b(\alpha_3 + \alpha_1 - \alpha_2)}.$$

where $\Upsilon_b(x)$ is the upsilon function and again take the large scaling limit $\alpha_i = \eta_i/b$, $b \rightarrow 0$ and $0 < \eta_i < \frac{1}{2}$ fixed for Seiberg bound.

- ▶ We can further divide the parameters into two classes:

$$\text{I: } \sum_i \eta_i > 1, \quad \text{II: } \sum_i \eta_i < 1, \quad \eta_i + \eta_j - \eta_k > 0.$$

Region I comes from convergence of the path integral, while Region II requires the complex saddles to make senses.

- ▶ For Region I, DOZZ coefficient reduces in this limit to:

$$C(\alpha_1, \alpha_2, \alpha_3) \sim \lambda^{(1-\sum_i \eta_i)/b^2} \exp \left[\frac{1}{b^2} \left\{ 1 - \sum_i \eta_i + F(2\eta_1) + F(2\eta_2) + F(2\eta_3) + F(0) \right. \right. \\ \left. \left. - F\left(\sum_i \eta_i - 1\right) - F(\eta_1 + \eta_2 - \eta_3) - F(\eta_2 + \eta_3 - \eta_1) - F(\eta_3 + \eta_1 - \eta_2) \right\} \right],$$

where $1/b^2$ is purely imaginary at leading order in $1/c$ expansion and $F(x)$ is a real function arising from the log $\Upsilon_b(x)$.

- ▶ The norm is thus:

$$| \langle V_{\alpha_1}(z_1, \bar{z}_1) V_{\alpha_2}(z_2, \bar{z}_2) V_{\alpha_3}(z_3, \bar{z}_3) \rangle |^2 \sim \mathcal{O}(1)$$

The obvious interpretation is that we cannot construct S^2 with three conical deficits given by $\eta_i \pi$ with $\sum_i \eta_i > 1$, hence no corresponding geometry.

- For Region II, using $\text{Re}[(\sum_i \eta_i - 1)/b^2] > 0$, we obtain:

$$\begin{aligned}
 C(\alpha_1, \alpha_2, \alpha_3) &\sim \left(e^{-\pi i \frac{1 - \sum_i \eta_i}{b^2}} - e^{\pi i \frac{1 - \sum_i \eta_i}{b^2}} \right) \lambda^{(1 - \sum_i \eta_i)/b^2} \\
 &\times \exp \left[\frac{1}{b^2} \left\{ F(2\eta_1) + F(2\eta_2) + F(2\eta_3) + F(0) - F\left(\sum_i \eta_i\right) \right. \right. \\
 &\quad - F(\eta_1 + \eta_2 - \eta_3) - F(\eta_2 + \eta_3 - \eta_1) - F(\eta_3 + \eta_1 - \eta_2) \\
 &\quad \left. \left. + 2 \left(1 - \sum_i \eta_i \right) \log \left(1 - \sum_i \eta_i \right) - 2 \left(1 - \sum_i \eta_i \right) \right\} \right].
 \end{aligned}$$

where again only two saddles contribute and its norm now yields:

$$| \langle V_{\alpha_1}(z_1, \bar{z}_1) V_{\alpha_2}(z_2, \bar{z}_2) V_{\alpha_3}(z_3, \bar{z}_3) \rangle |^2 \sim \exp \left[\frac{\pi c^{(g)}}{3} \left(1 - \sum_i \eta_i \right) \right]$$

Most of the exponents cancel out due to the overall purely imaginary factor $\frac{1}{b^2}$.

- ▶ We propose the bulk geometry to be:

$$ds^2 = d\theta^2 + \cos^2 \theta ds_{\text{con}}^2$$

where ds_{con}^2 denotes the metric of S^2 with three conical deficits [Umehara, Yamada 2000], where each deficit is created by the heavy vertex operator insertion with deficit angle $4\pi\eta_i$.

- ▶ The resultant volume is $(1 - \sum_i \eta_i) \geq 0$ fraction of S^3 , reproducing the results from Liouville three point function in this limit.

So far we have only discussed the dS case, the situation with AdS is also interesting and somewhat puzzling.

- ▶ We can see this from the exact two point function but now with $\text{Re}(b^{-2}) > 0$, the analytic continuation of Γ -function yields:

$$\langle V_\alpha(z_1)V_\alpha(z_2) \rangle \sim \frac{1}{e^{-\frac{\pi i}{b^2}(1-2\eta)} - e^{\frac{\pi i}{b^2}(1-2\eta)}} |z_{12}|^{-\frac{4}{b^2}\eta(1-\eta)} e^{-\frac{2}{b^2}[(1-2\eta)\ln(1-2\eta)-(1-2\eta)]}$$

after the expanding the denominator, it becomes an infinite series:

$$\frac{1}{e^{-\frac{\pi i}{b^2}(1-2\eta)} - e^{\frac{\pi i}{b^2}(1-2\eta)}} \sim e^{\pi i(1-2\eta)\frac{\ell_{\text{AdS}}}{4G}} \sum_{n=0}^{\infty} e^{n\pi i(1-2\eta)\frac{\ell_{\text{AdS}}}{2G}}.$$

- ▶ We can use holography dictionary to relate that $\mathcal{Z}_{\text{AdS}} = \langle V_\alpha(z_1)V_\alpha(z_2) \rangle$, we have infinite many saddles:

$$\mathcal{Z}_{\text{AdS}} = \sum_{n=0}^{\infty} \mathcal{Z}_n$$

$$\mathcal{Z}_n \sim e^{\frac{\ell_{\text{AdS}}}{4G}(2n+1)\pi i(1-2\eta)} |z_{12}|^{-\frac{\ell_{\text{AdS}}}{2G}\eta(1-\eta)} e^{-\frac{\ell_{\text{AdS}}}{2G}[(1-2\eta)\ln(1-2\eta)-(1-2\eta)]}.$$

- ▶ It is interesting to consider the complexification of euclidean AdS_3 :

$$ds^2 = \ell_{\text{AdS}}^2 \left[\left(\frac{d\theta(u)}{du} \right)^2 du^2 + \sinh^2 \theta(u) d\Sigma^2 \right]$$

where $\theta(u)$ is a holomorphic function of u . If we consider geometries approach to AdS_3 as $u \rightarrow \infty$ and truncates at $u = 0$, thus need $\theta \rightarrow u, u \rightarrow \infty$ and $u = in\pi, u = 0$.

- ▶ The two geometries may be interpolated via:

$$\theta = n\pi i(1 - u) \quad (0 \leq u \leq 1), \quad \theta = (u - 1) \quad (u > 1)$$

while this yields EAdS_3 for $u > 1$, for $0 \leq u \leq 1$ the geometry becomes multiple wrapping over S^3 of imaginary radius $i\ell_{\text{AdS}}$.

- ▶ This may seem somewhat unphysical, even though from Chern-Simons gauge theory, we can construct configuration with action $2\pi in$, where n labels π_3 of imaginary S^3 .

We can also consider such interpolation in the embedding space.

- ▶ Starting with the asymptotic geometry of Euclidean AdS₃:

$$\tilde{X}_0^2 + X_1^2 + X_2^2 - X_3^2 = -\ell_{\text{AdS}}^2$$

We may want to interpolate it to Lorentzian AdS₃ given by:

$$-X_0^2 + X_1^2 + X_2^2 - X_3^2 = -\ell_{\text{AdS}}^2$$

by setting $i\tilde{X}_0 = X_0$, however Lorentzian AdS₃ has trivial topology.

- ▶ Instead we may consider the interpolation $X_3 = i\tilde{X}_3$, such that:

$$\tilde{X}_0^2 + X_1^2 + X_2^2 + \tilde{X}_3^2 = -\ell_{\text{AdS}}^2$$

with $|X_3| \geq \ell_{\text{AdS}}$. We can glue the two geometries at $X_3 = i\tilde{X}_3 = \ell_{\text{AdS}}$ and $X_1 = X_2 = \tilde{X}_0 = 0$.

Another interesting approach to explore the interpolation is mini-superspace approach. [Chen, Hikida, Taki, Uetoko, 2024]

- ▶ Starting with Einstein-Hilbert action for positive cosmological constant and contact term:

$$I = -\frac{1}{16\pi G} \int d^3x \sqrt{g} \left(R - \frac{2}{\ell_{\text{dS}}^2} \right) + I_{\text{bdy}}. \quad I_{\text{CT}} = \frac{1}{8\pi G} \int d^2x \sqrt{h} \sqrt{-\frac{1}{\ell_{\text{dS}}^2}}$$

- ▶ Substituting the homogeneous metric ansatz:

$$ds^2 = \ell_{\text{dS}}^2 [N(\tau)^2 d\tau^2 + a(\tau)^2 d\Omega_2]$$

where $0 \leq \tau \leq 1$ and we can use the gauge redundancy to fix $N(\tau) = N$. The gravitational path integral is now reduced to:

$$\Psi = \int_{\mathcal{C}} dN \int \mathcal{D}a e^{-I[a;N] - I_{\text{CT}}}$$
$$I[a; N] = -\frac{\ell_{\text{dS}}}{2G} \int_0^1 d\tau N \left(\frac{1}{N^2} \left(\frac{da}{d\tau} \right)^2 - a^2 + 1 \right) + (\text{boundary contributions})$$

We thus need to evaluate its saddle contributions and consider \mathcal{C} .

We reduce the problem into solving for $a(\tau)$ and fluctuations around it.

- ▶ If we impose the Dirichlet boundary conditions $a(0) = 0$ and $a(1) = a_1 > 0$, the solution to E.O.M for $a(\tau)$ is:

$$\bar{a}^{(N)}(\tau) = \frac{a_1}{\sin N} \sin(N\tau)$$

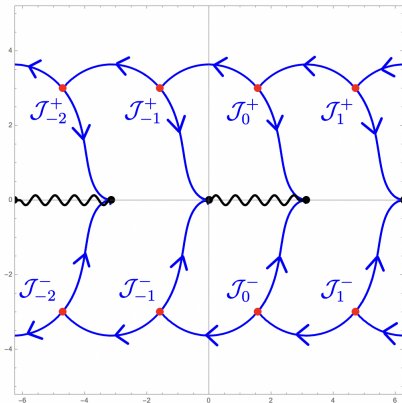
If we include the fluctuations around $\bar{a}^{(N)}(\tau)$, the path integral for N becomes:

$$\Psi = \int_{\mathcal{C}} dN \left(\frac{1}{\sqrt{N} \sin N} \right)^{\frac{1}{2}} e^{-I[\bar{a}^{(N)}; N] - I_{CT}}$$
$$I[\bar{a}^{(N)}; N] = -\frac{\ell_{dS}}{2G} (N + a_1^2 \cot N)$$

- ▶ To work out \mathcal{C} , we first consider the saddle for N satisfying $\partial I[\hat{a}, N]/\partial N = 0$, the solutions are ($m \in \mathbb{Z}$):

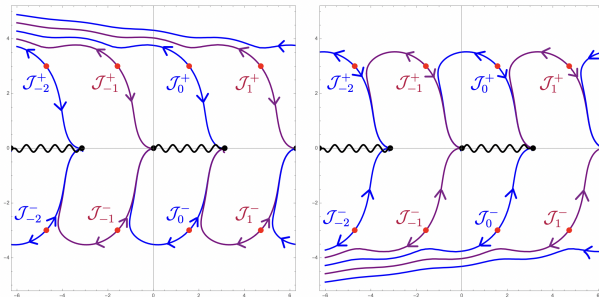
$$N_m^+ = \left(m + \frac{1}{2}\right) \pi + i \log \left(a_1 + \sqrt{a_1^2 - 1}\right)$$
$$N_m^- = \left(m + \frac{1}{2}\right) \pi - i \log \left(a_1 + \sqrt{a_1^2 - 1}\right)$$

- ▶ It is interesting to note that in complex N -plane, the steepest descent emanating from one saddle point can end up at another saddle point (red dots):



This implies that we have Stoke's phenomenon, and makes the identification of \mathcal{C} difficult.

- ▶ Instead we can consider slight deformation of contours by deforming $l_{\text{dS}} \rightarrow l_{\text{dS}} \pm i\epsilon$:



We can identify the correct shift by interpreting it as subleading correction in:

$$c = -i \frac{3l_{\text{dS}}}{2G} + 13 + \mathcal{O}(G).$$

such that it should be $l_{\text{dS}} \rightarrow l_{\text{dS}} + i\epsilon$.

- ▶ To complete the construction of \mathcal{C} , notice that each saddle N_m^\pm :

$$\Psi_m^\pm \sim e^{\frac{(2m+1)\ell_{\text{dS}}\pi}{4G}} (2a_1)^{\mp i \frac{\ell_{\text{dS}}}{2G} \pm \frac{\epsilon}{2G}} .$$

where in large a_1 limit, Ψ_m^- in the lower half becomes suppressed. Since N acts as time-direction, naively we should integrate along $i\mathbb{R}$ in the upper half or deforming it into \mathcal{J}_{-1}^+ , however it implies only exponentially suppressed saddle N_{-1}^+ which is insufficient.

- ▶ This implies that we should also pick up additional saddles N_0^+ and the contour which does this can be deformed into:

$$-\mathcal{J}_{-1}^+ + \mathcal{J}_0^- + \mathcal{J}_0^+ ,$$

which goes around the branch cut and back to upper half plane. The contribution from N_0^- in lower half disappears in large a_1 limit.

We perform similar analysis for the negative cosmological constant case, here are the summary.

- ▶ The solution for Dirichlet boundary conditions yields:

$$\bar{a}^{(N)}(r) = \frac{a_1}{\sinh N} \sinh(Nr).$$

where the fluctuation determinant around it can be evaluated analogous to yield:

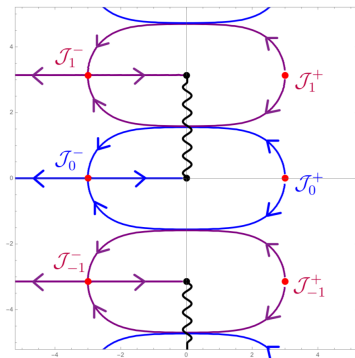
$$\mathcal{Z} = \int_{\mathcal{C}} dN \left(\frac{1}{\sqrt{N} \sinh N} \right)^{\frac{1}{2}} e^{-I[\bar{a}^{(N)}; N] - I_{\text{CT}}}$$

$$I[\bar{a}^{(N)}; N] = -\frac{\ell_{\text{AdS}}}{2G} [N + a_1^2 \coth N]$$

- ▶ The one-loop corrected action yields the following saddle points:

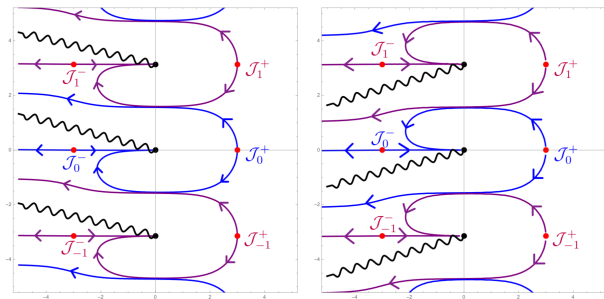
$$\begin{aligned} N_m^+ &= \operatorname{arcsinh} a_1 + \pi i m \quad (m \in \mathbb{Z}), \\ N_m^- &= -\operatorname{arcsinh} a_1 + \pi i m \quad (m \in \mathbb{Z}), \end{aligned}$$

The saddle points and branch cuts are given here:



Here we again see that there are Stoke's phenomena where paths of steepest descent connecting saddle points, but moreover we have the thimble \mathcal{J}_m^+ crossing the branch cut. These make the identification of \mathcal{C} difficult.

If we now deform the contour by shifting $\ell_{\text{AdS}} \rightarrow \ell_{\text{AdS}} \pm i\epsilon$, and choose the branch cuts so as to avoid the steepest descent paths, we obtain:



Here the individual saddle contribute as:

$$\mathcal{Z}_m^\pm \sim e^{\frac{im\pi\ell_{\text{AdS}}}{2G}} (2a_1)^\pm \frac{\ell_{\text{AdS}}}{2G}$$

for saddle labeled by N_m^\pm respectively.

- ▶ For $\ell_{\text{AdS}} \rightarrow \ell_{\text{AdS}} + i\epsilon$, the final total contour is:

$$\mathcal{C} \rightarrow \sum_{m=0}^{\infty} \mathcal{J}_m^+ - \sum_{m=1}^{\infty} \mathcal{J}_m^- ,$$

which yields:

$$\mathcal{Z} \sim \sum_{m=0}^{\infty} e^{\frac{i\pi m(\ell_{\text{AdS}} + i\epsilon)}{2G}} (2a_1)^{\frac{\ell_{\text{AdS}} + i\epsilon}{2G}} = -\frac{2ie^{-\frac{i\pi\ell_{\text{AdS}}}{4G}}}{\sin\left(\frac{\ell_{\text{AdS}}\pi}{4G}\right)} (2a_1)^{\frac{\ell_{\text{AdS}}}{2G}} .$$

which differs from the CFT result by an overall phase factor.

- ▶ For $\ell_{\text{AdS}} \rightarrow \ell_{\text{AdS}} - i\epsilon$, the final total contour is:

$$\mathcal{C} \rightarrow -\sum_{m=1}^{\infty} \mathcal{J}_{-m}^+ + \sum_{m=0}^{\infty} \mathcal{J}_{-m}^- .$$

which yields:

$$\mathcal{Z} \sim -\sum_{m=-\infty}^0 e^{\frac{i\pi m(\ell_{\text{AdS}} - i\epsilon)}{2G}} (2a_1)^{\frac{\ell_{\text{AdS}} - i\epsilon}{2G}} = -\frac{2ie^{\frac{i\pi\ell_{\text{AdS}}}{4G}}}{\sin\left(\frac{\ell_{\text{AdS}}\pi}{4G}\right)} (2a_1)^{\frac{\ell_{\text{AdS}}}{2G}} .$$

which again differs from the CFT result by an overall phase factor.

The analysis allows us to simply pick the contour \mathcal{C} to be \mathbb{R}_+

- ▶ In this talk we demonstrate how holography may help us to explore the complex geometries arising from the gravitational path integral.
- ▶ We actually also studied higher point CFT correlation functions with heavy operator insertions which can be dual to Chern-Simons Wilson lines and compute their monodromy matrices.
- ▶ It would be interesting to extend the analysis presented to higher point CFT correlation functions and CFT on higher genus Riemann surfaces.
- ▶ Extensions to higher spin gravity theories and the dual $SL(N)$ Toda CFT may also be interesting.